



TENSION WITH THE Λ CDM MODEL *AT HIGH REDSHIFT FROM A HUBBLE DIAGRAM OF QUASARS*

Elisabetta Lusso

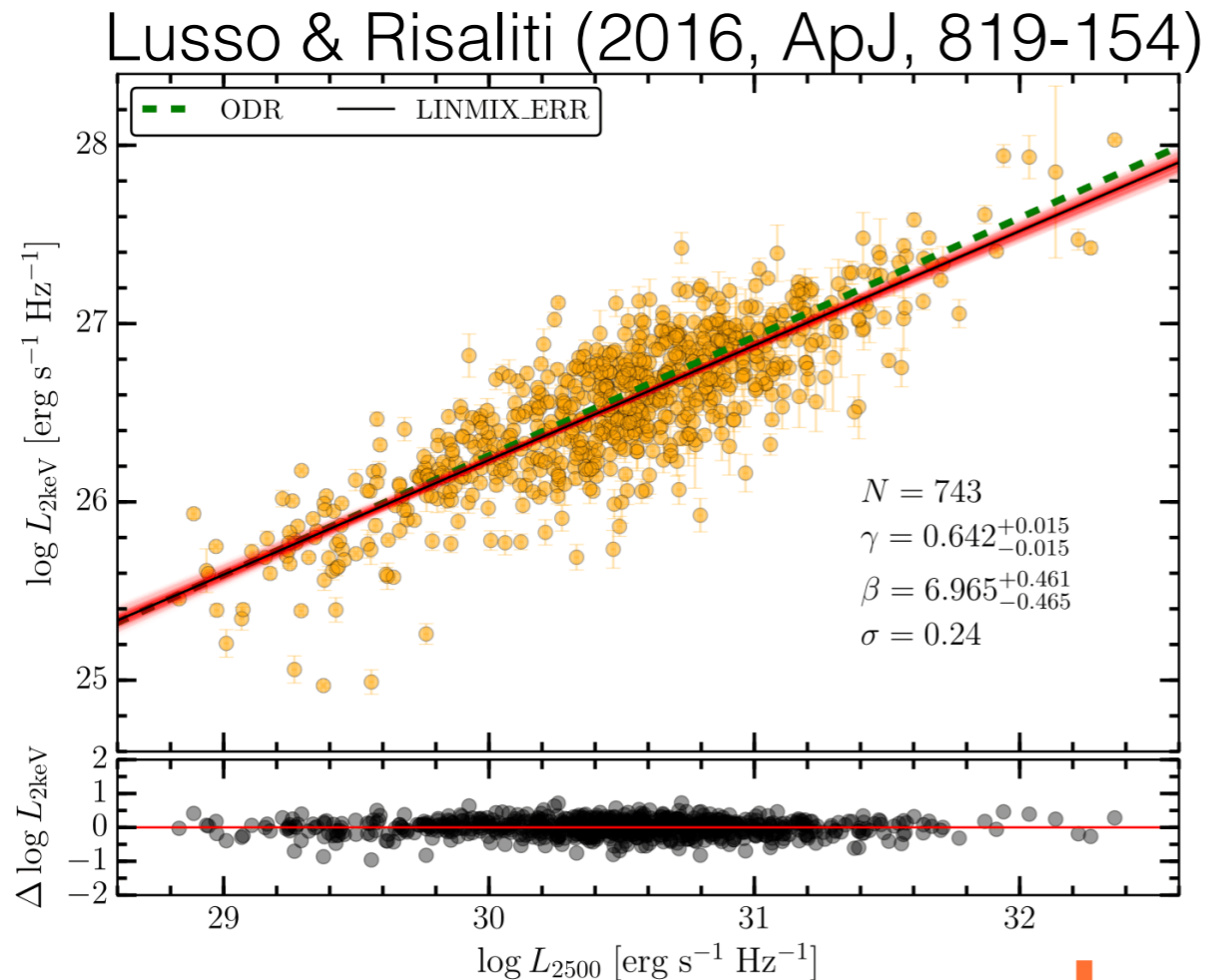
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Francesco Salvestrini, Francesca Civano, Ester Piedipalumbo,
Maurizio Paolillo, Lorenzo Amati*

*Supermassive Black holes: evolution & environment
Corfu 19-22 June*

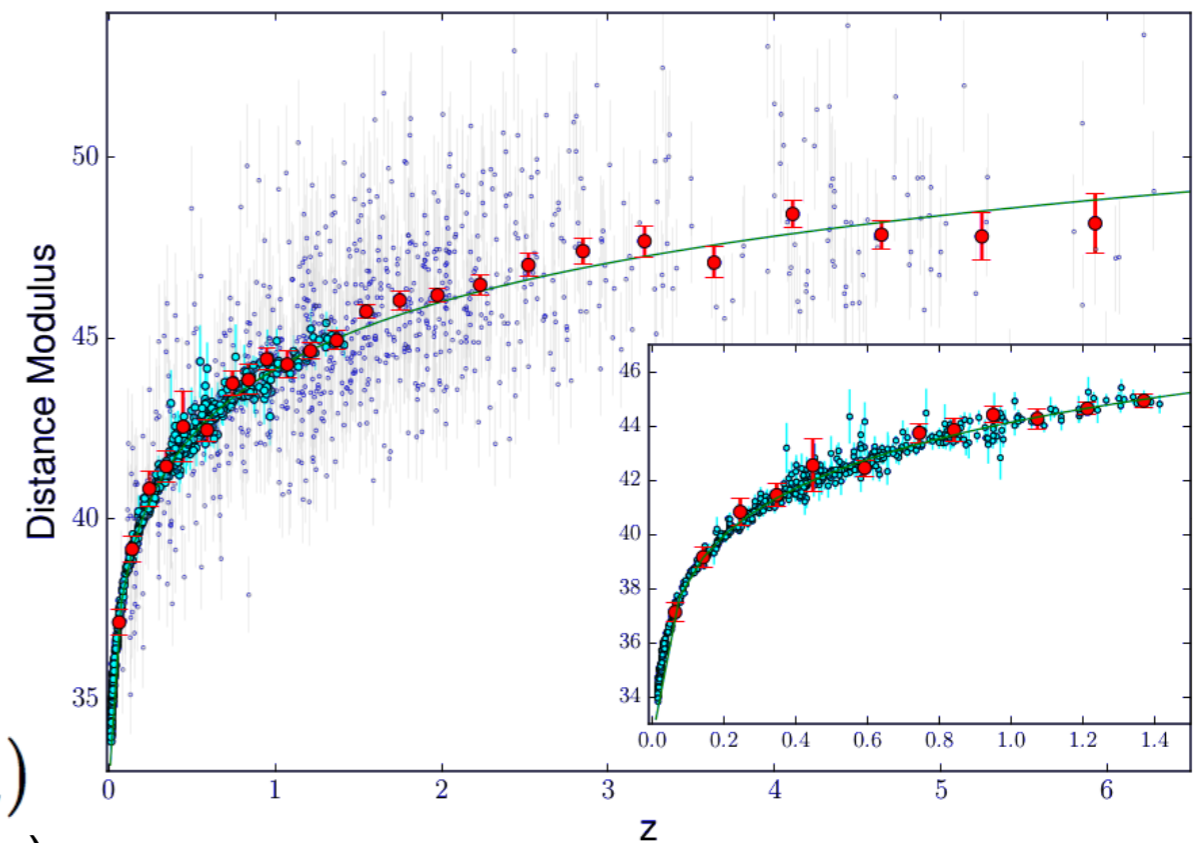
Cosmology with quasars

The distance modulus- z relation



$$\log(L_X) = \beta + \gamma \log(L_{UV})$$

Risaliti & Lusso (2015, ApJ, 815-33)



Standardise the quasar emission

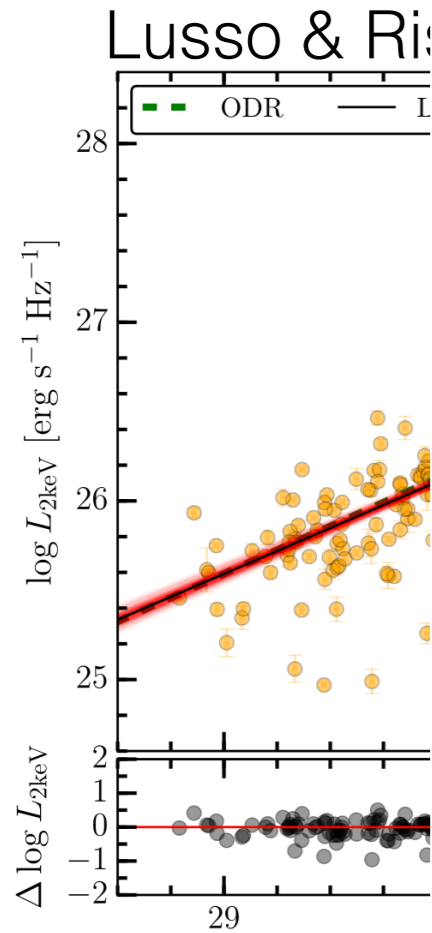
$$\begin{aligned} \log(F_X) &= \Phi(F_{UV}, D_L) \\ &= \beta' + \gamma \log(F_{UV}) + 2(\gamma - 1)\log(D_L) \\ &\quad D_L(z, \Omega_M, \Omega_\Lambda) \end{aligned}$$

The L_X - L_{UV} non-linear relation as a way to measure quasar distances

Cosmology with quasars

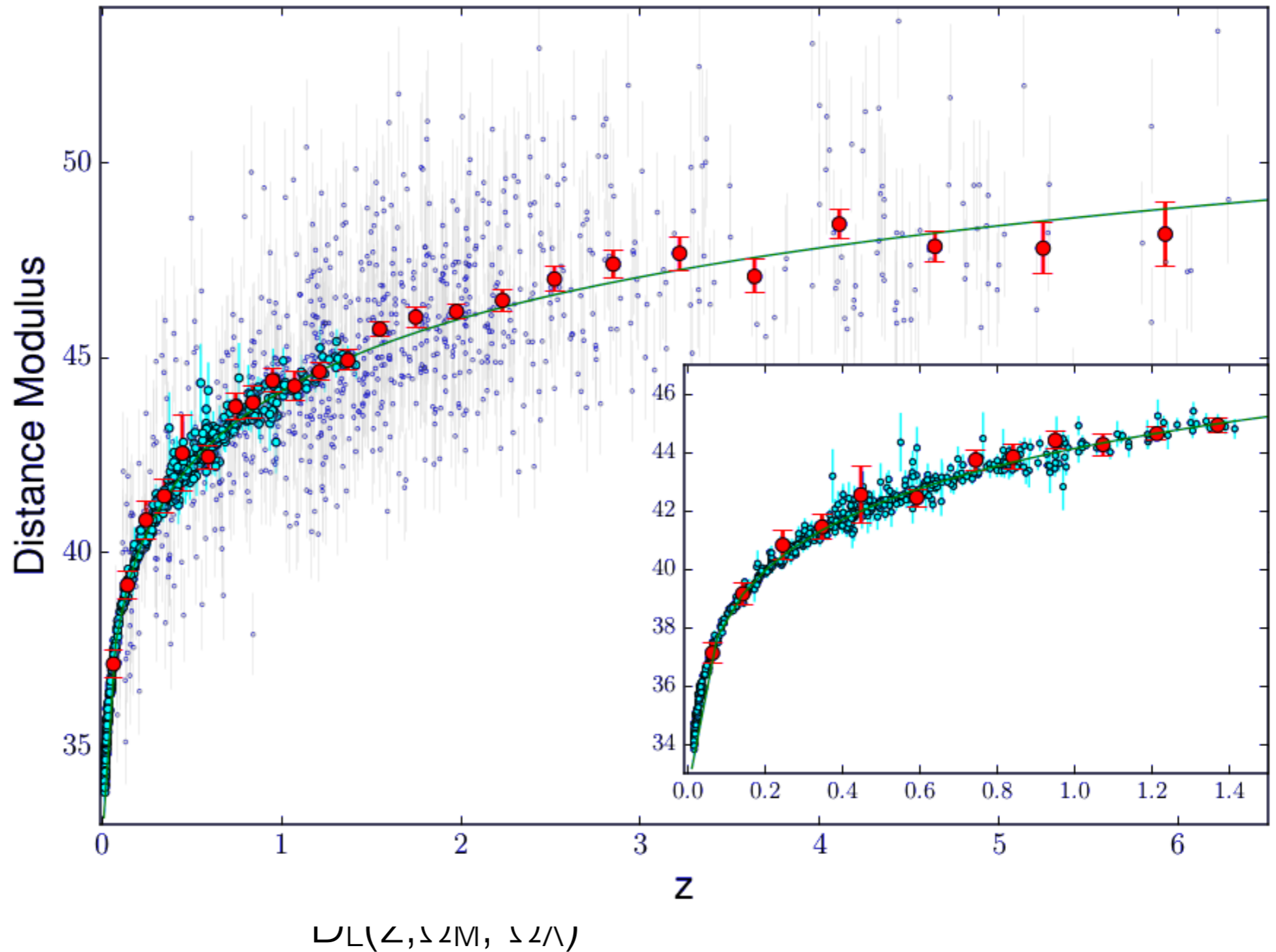
Risaliti & Lusso (2015, ApJ, 815-33)

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Standardise

$$\log(F_X) = \Phi(z) + \beta' \log(L_{UV})$$

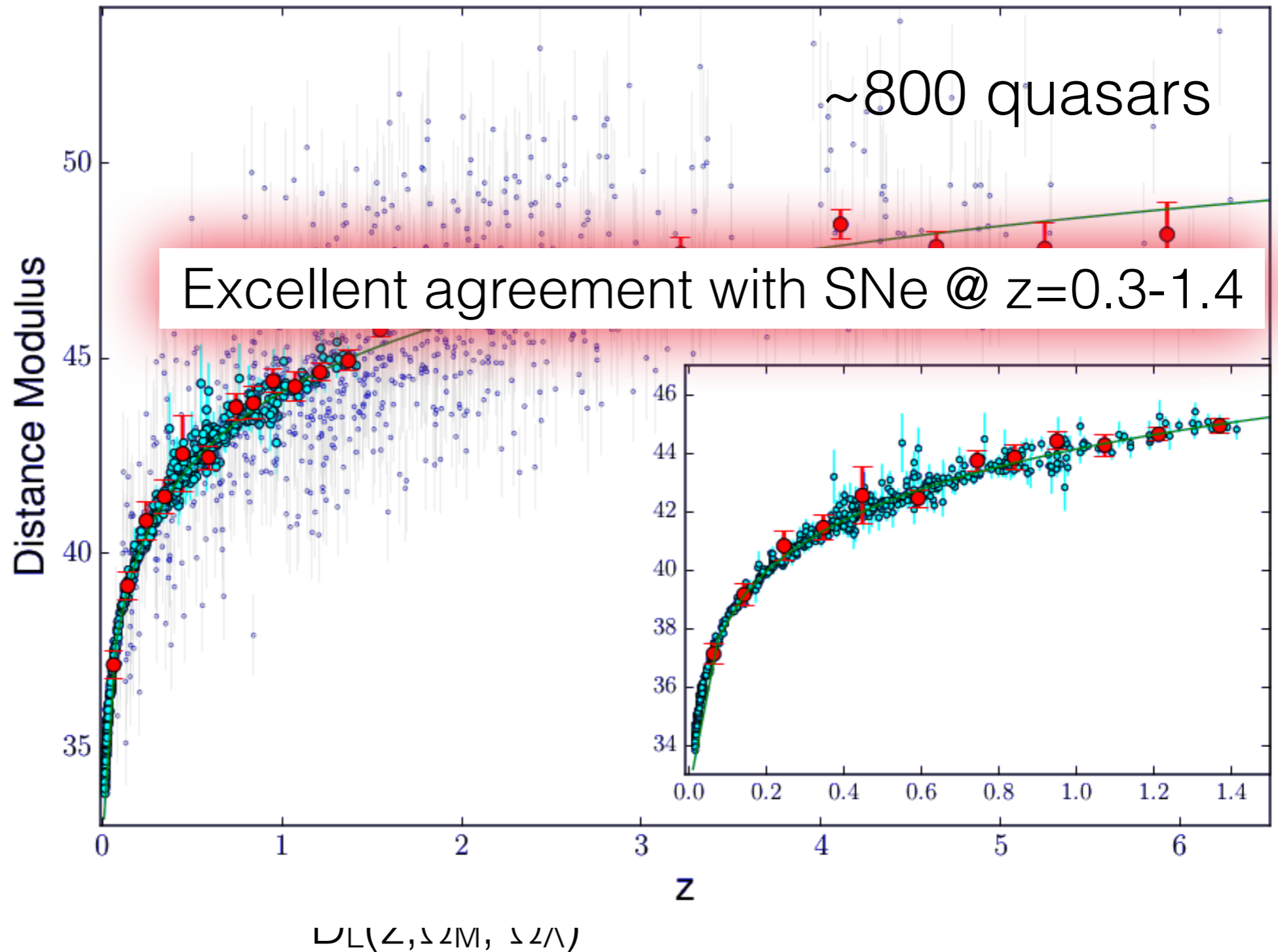
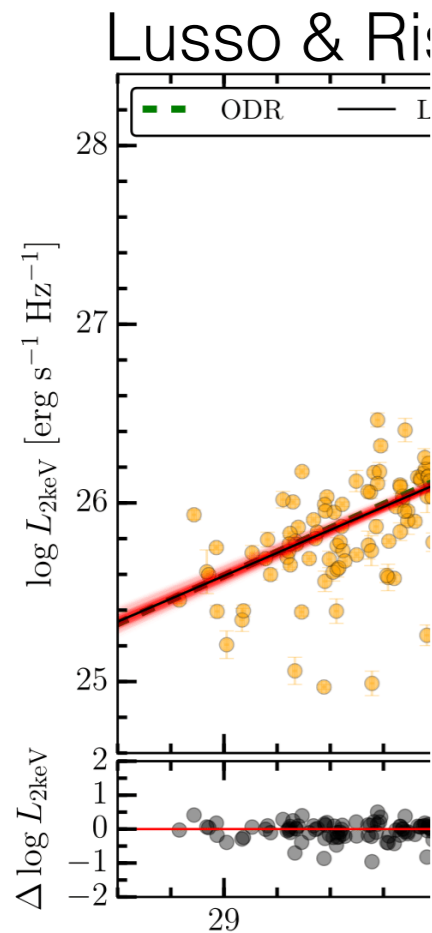


The L_X - L_{UV} non-linear relation as a way to measure quasar distances

Cosmology with quasars

Risaliti & Lusso (2015, ApJ, 815-33)

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Standardise

$$\log(F_X) = \Phi(\dots) = \beta'$$

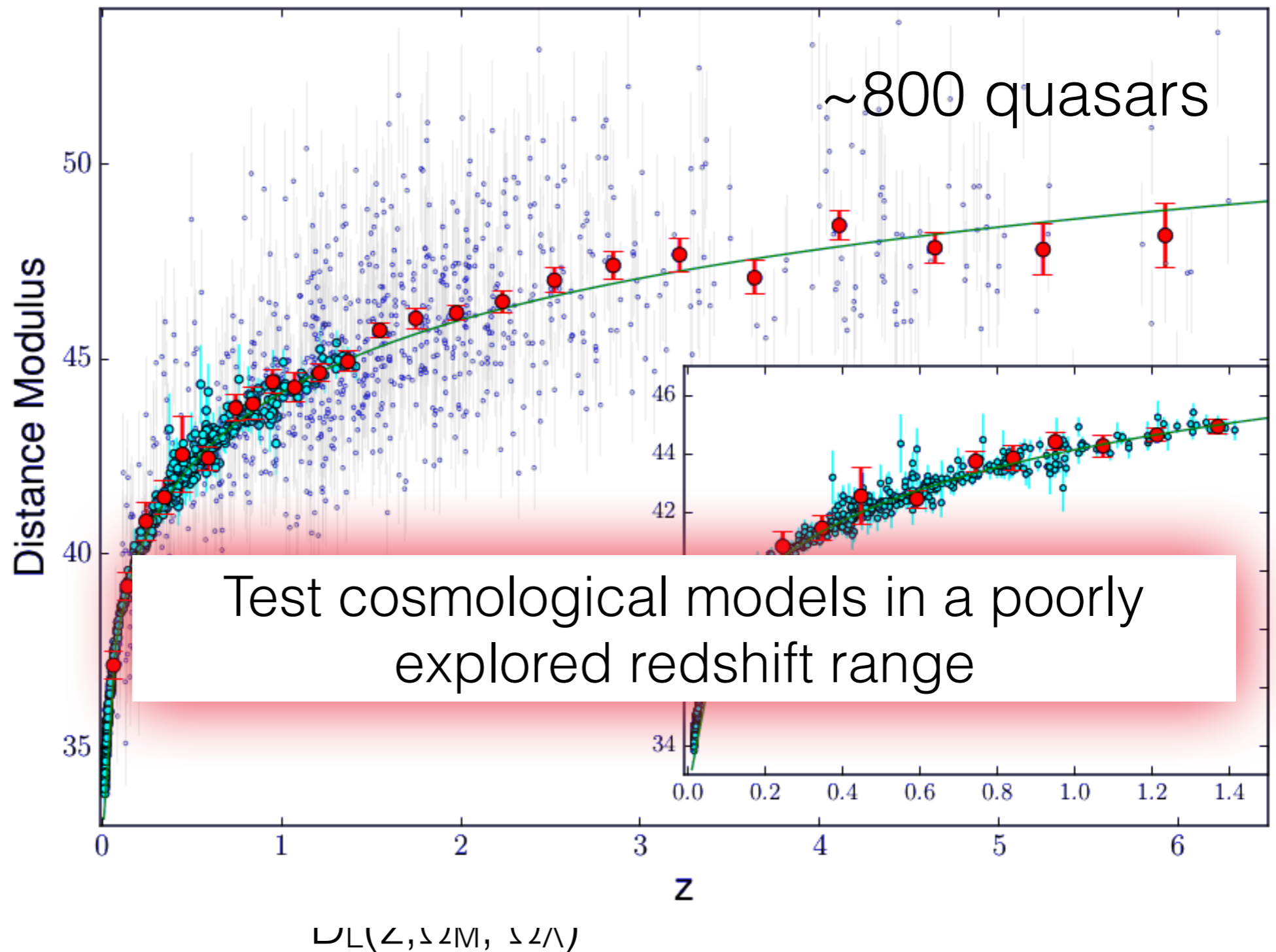
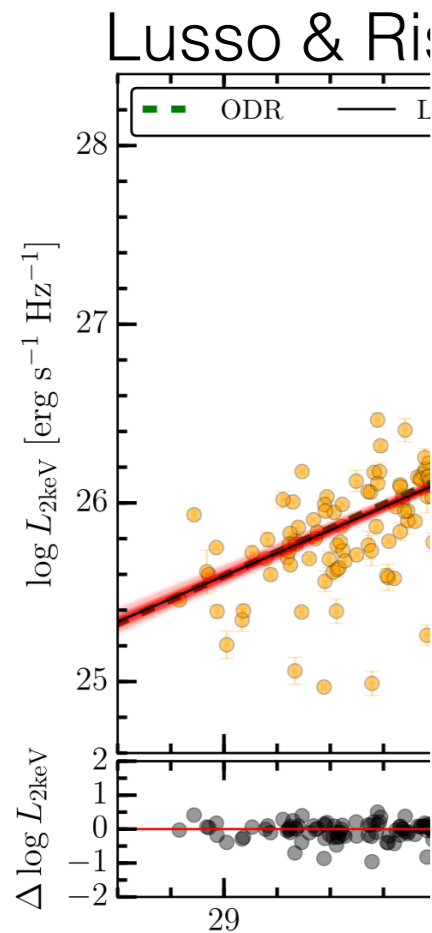
$\log(L(z, \Delta z_M, \Delta z_\Lambda))$

The L_X - L_{UV} non-linear relation as a way to measure quasar distances

Cosmology with quasars

Risaliti & Lusso (2015, ApJ, 815-33)

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Standardise

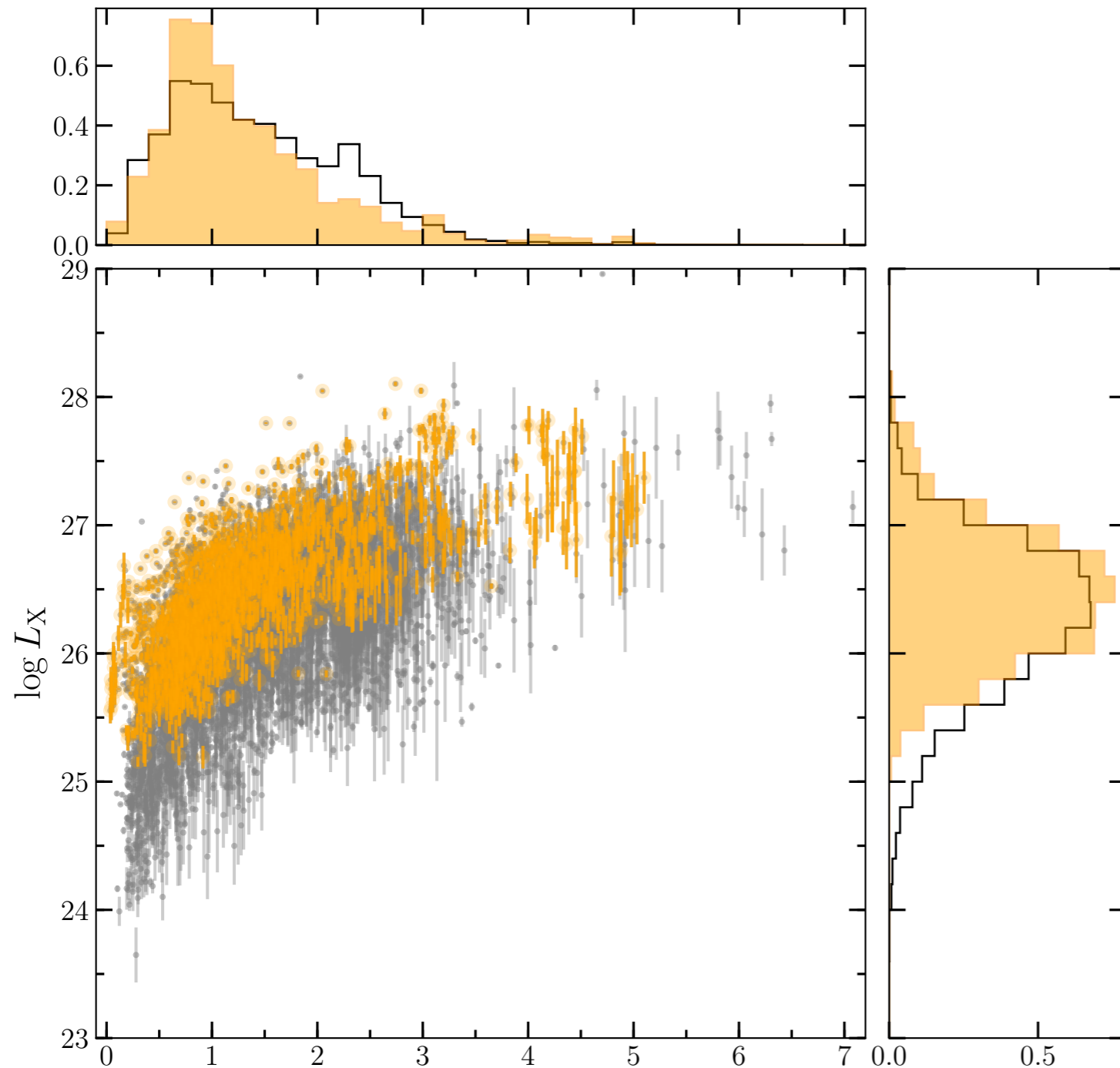
$$\log(F_X) = \Phi(\dots) = \beta'$$

The L_X - L_{UV} non-linear relation as a way to measure quasar distances

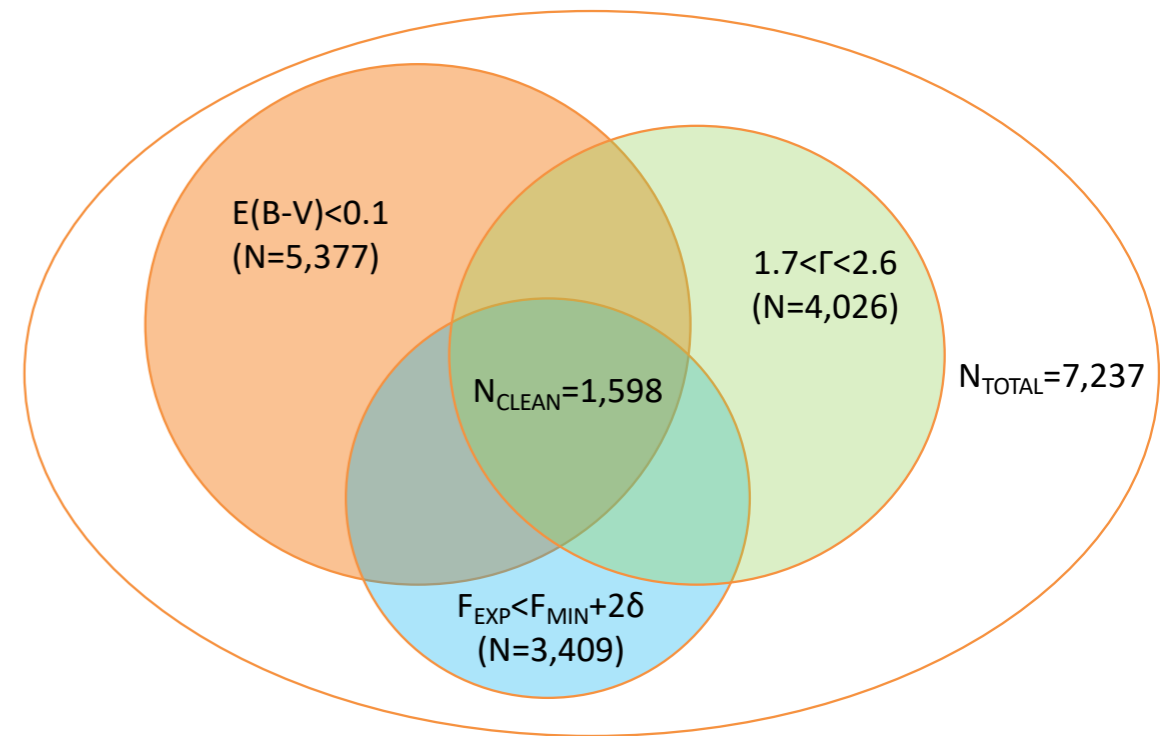
Cosmology with quasars

The new! Quasars Hubble Diagram: sample

~1600 quasars: SDSS+3XMM+XMM LP+archive/literature



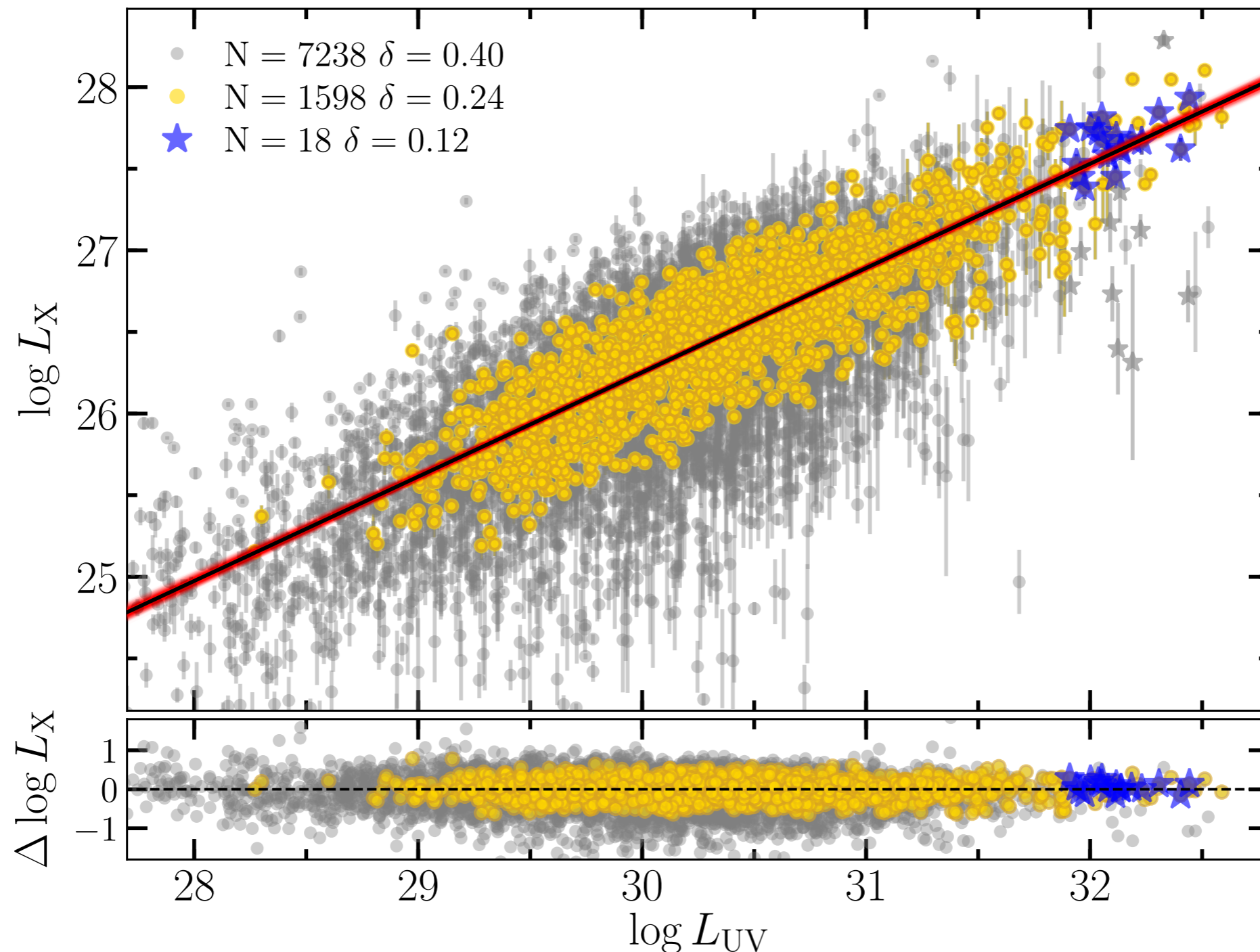
Risaliti & Lusso (2019) z



Cosmology with quasars

The new! Quasars Hubble Diagram: L_X - L_{UV}

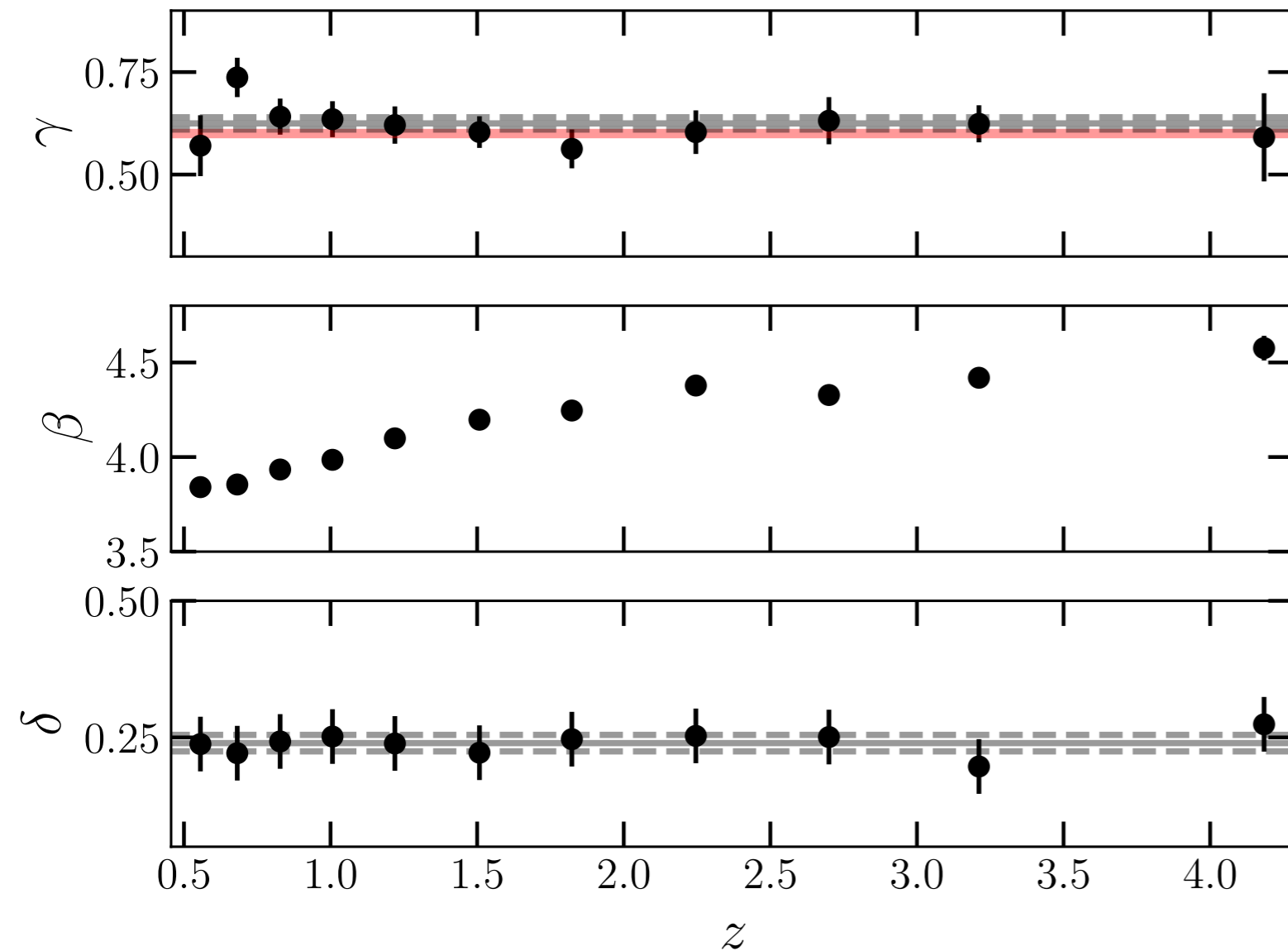
~1600 quasars: SDSS+3XMM+XMM LP+archive/literature



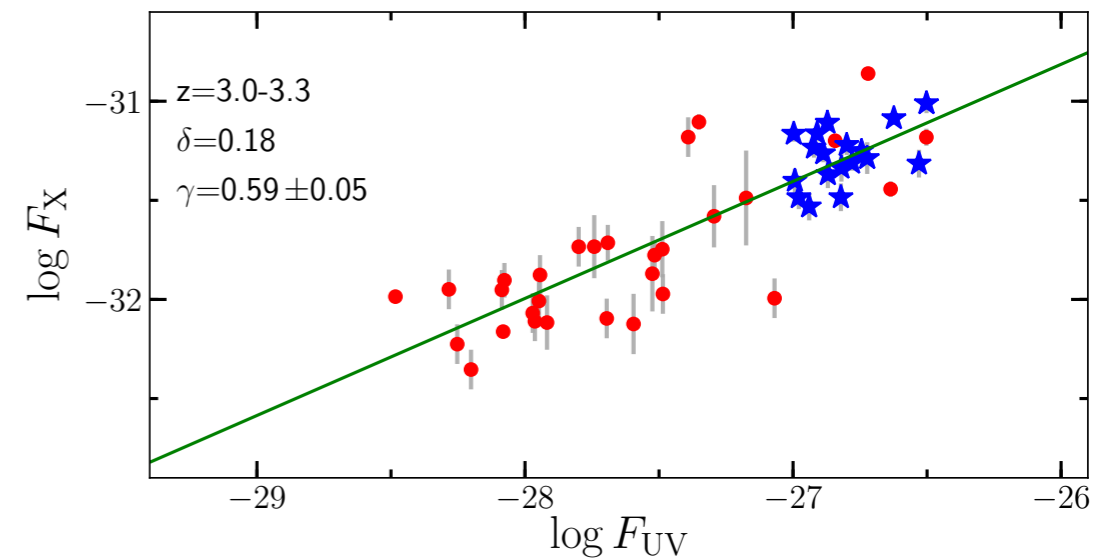
Risaliti & Lusso (2019)

Cosmology with quasars

The new! Quasars Hubble Diagram: redshift dependence



$$\log D_L = \frac{1}{2-2\gamma} (\gamma \log F_{UV} - \log F_X) + \beta.$$

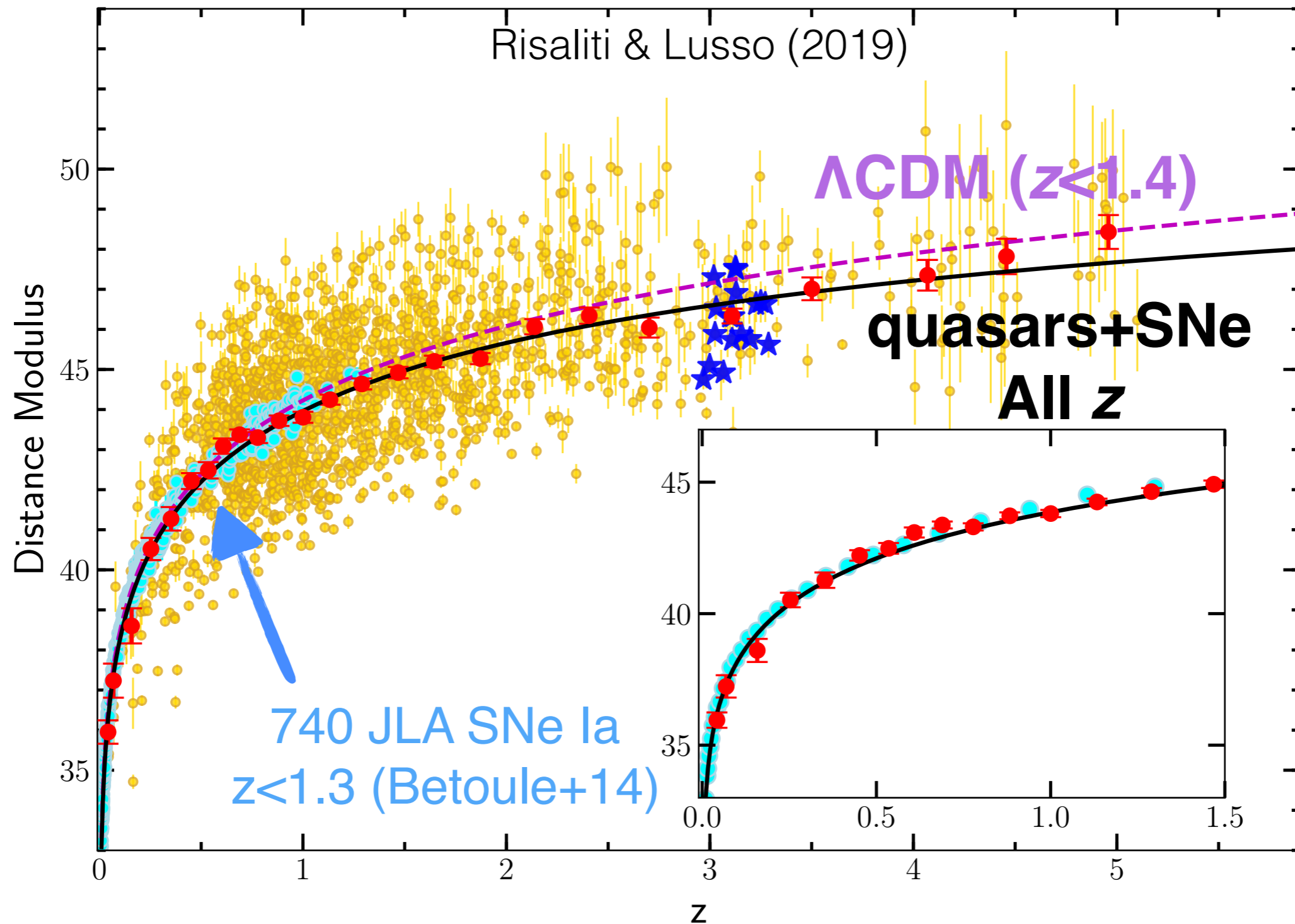


Risaliti & Lusso (2019)

Cosmology with quasars

The new! Quasars Hubble Diagram

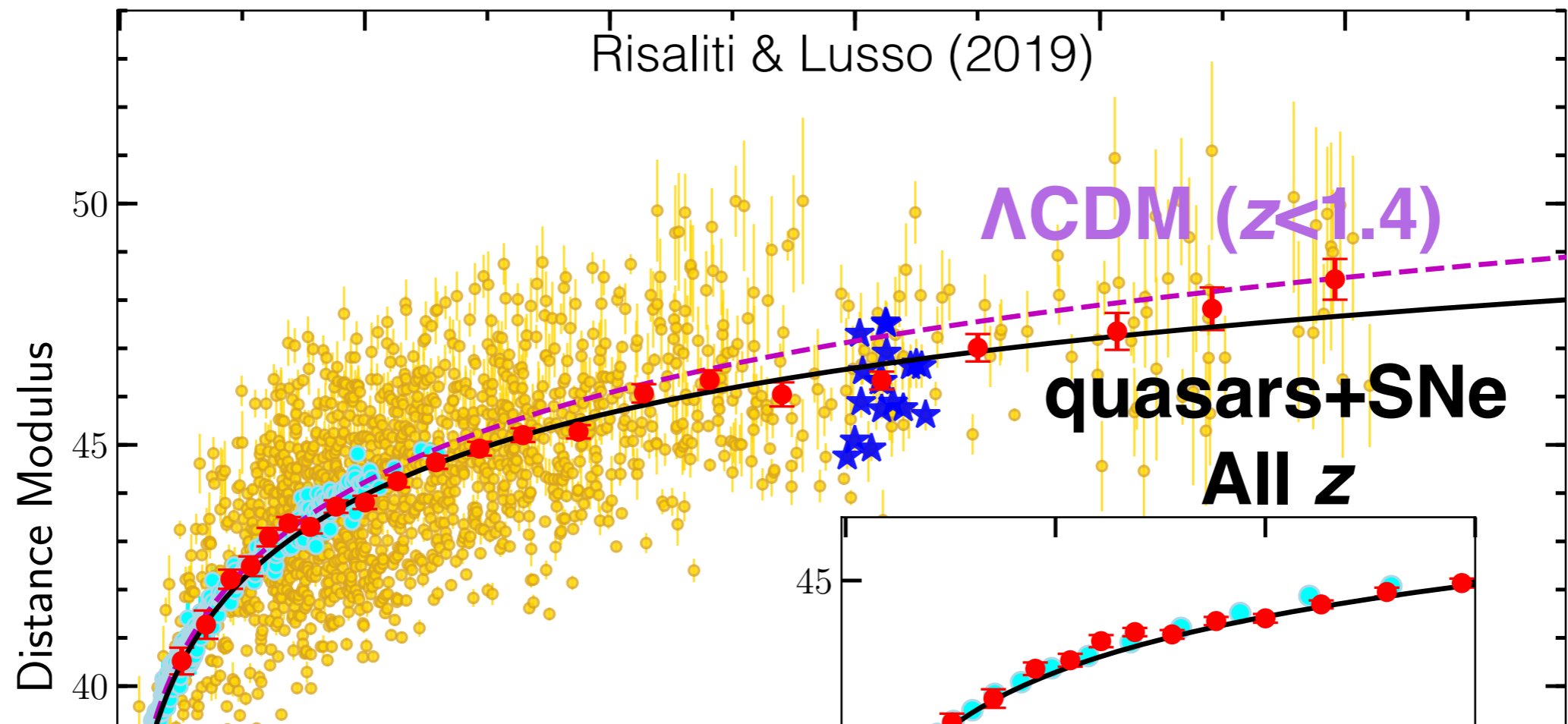
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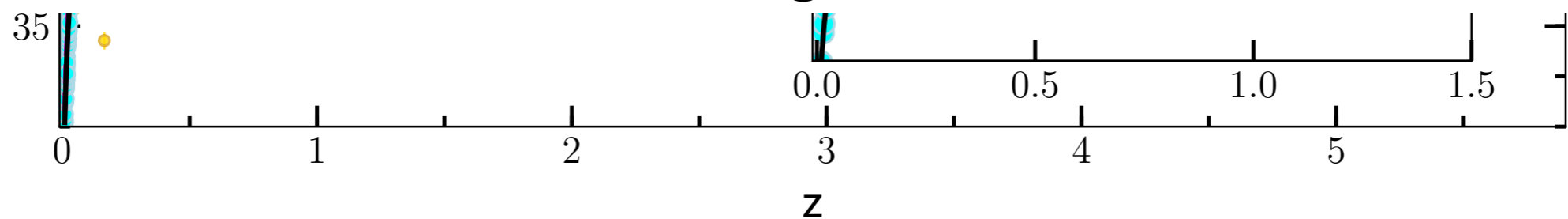
Cosmology with quasars

The new! Quasars Hubble Diagram

~1600 quasars: SDSS+3XMM+XMM LP+archive/literature



deviation from the Λ CDM model emerges at higher redshift,
with a statistical significance of $\sim 4\sigma$



Cosmology with quasars

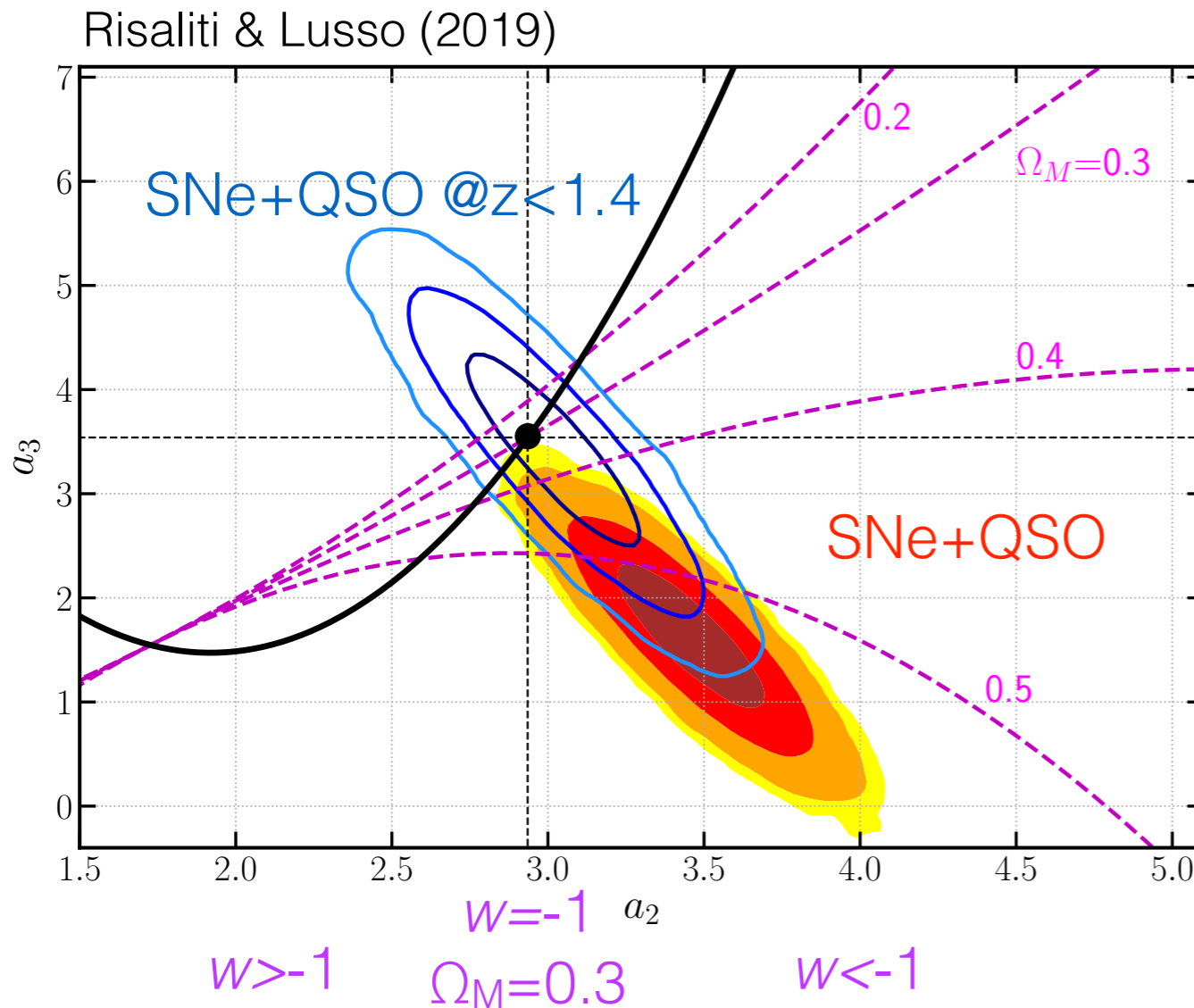
The new! Quasars Hubble Diagram

Cosmographic approach

$$P[\log(1+z)]: D_L = k \sum_i a_i [\log(1+z)]^i$$

$$k = \ln(10)c/H_0$$

$$a_2(\Omega_M), a_3(\Omega_M)$$



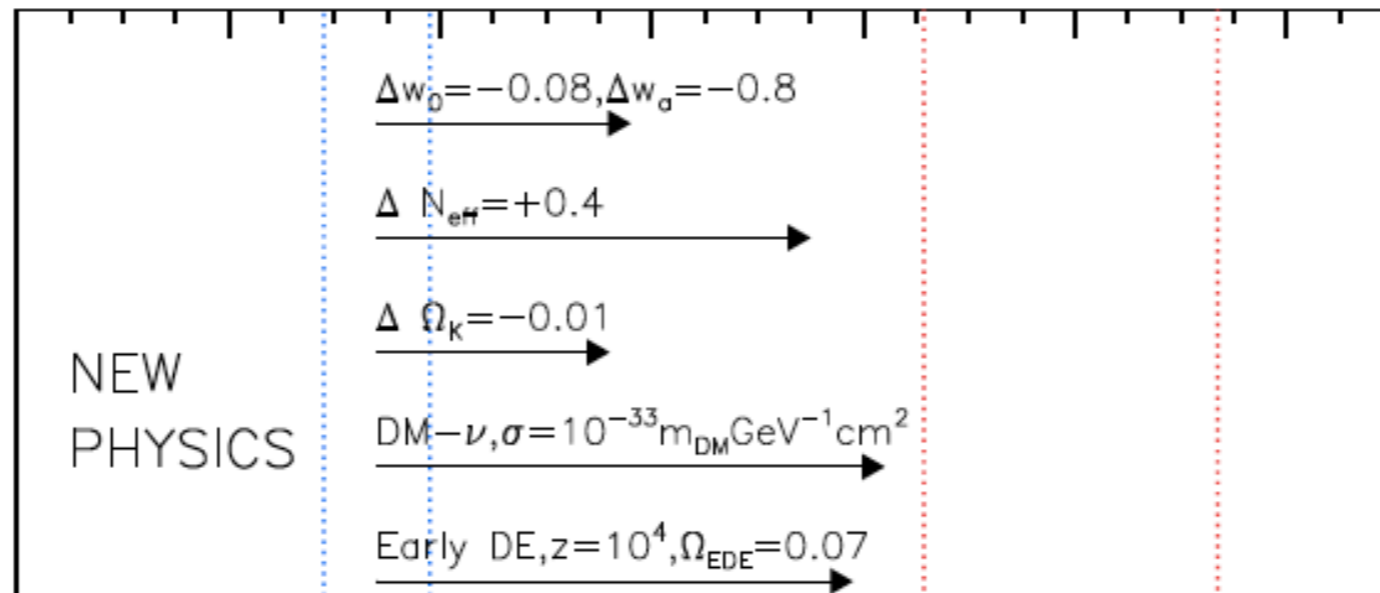
- Flat Λ CDM
- - - Flat w CDM (free w)
 $w(z) = w_0 + w_a \cdot z / (1+z)$

data suggest: **dark energy density increasing with time.**

Within the w CDM model: $\Omega_M > 0.3$ and $w < -1.3$

The schism between the early & late Universe

Do we need an extension to the Λ CDM? Maybe yes...

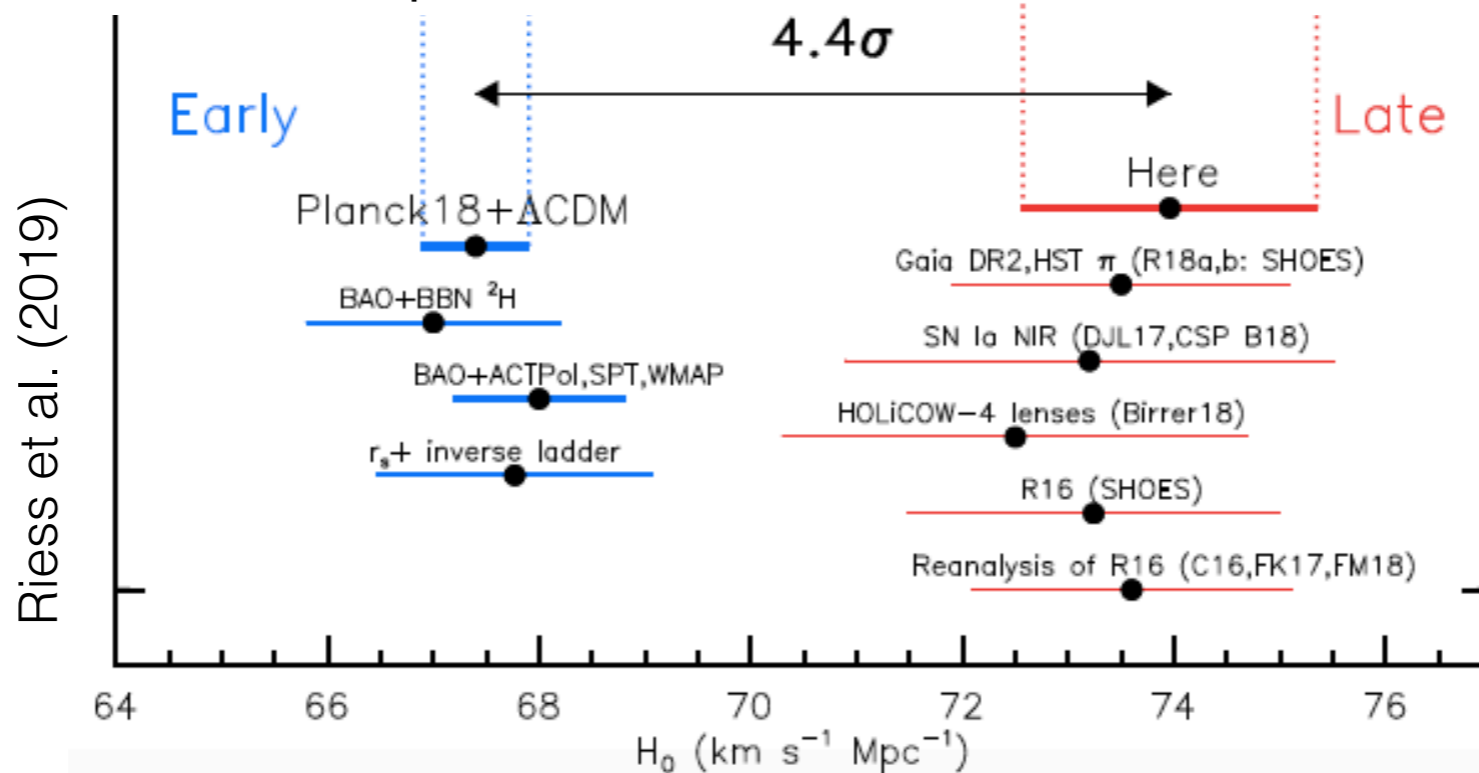


Planck et al. (2018)

$$H_0 = 67.06 \pm 0.62 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

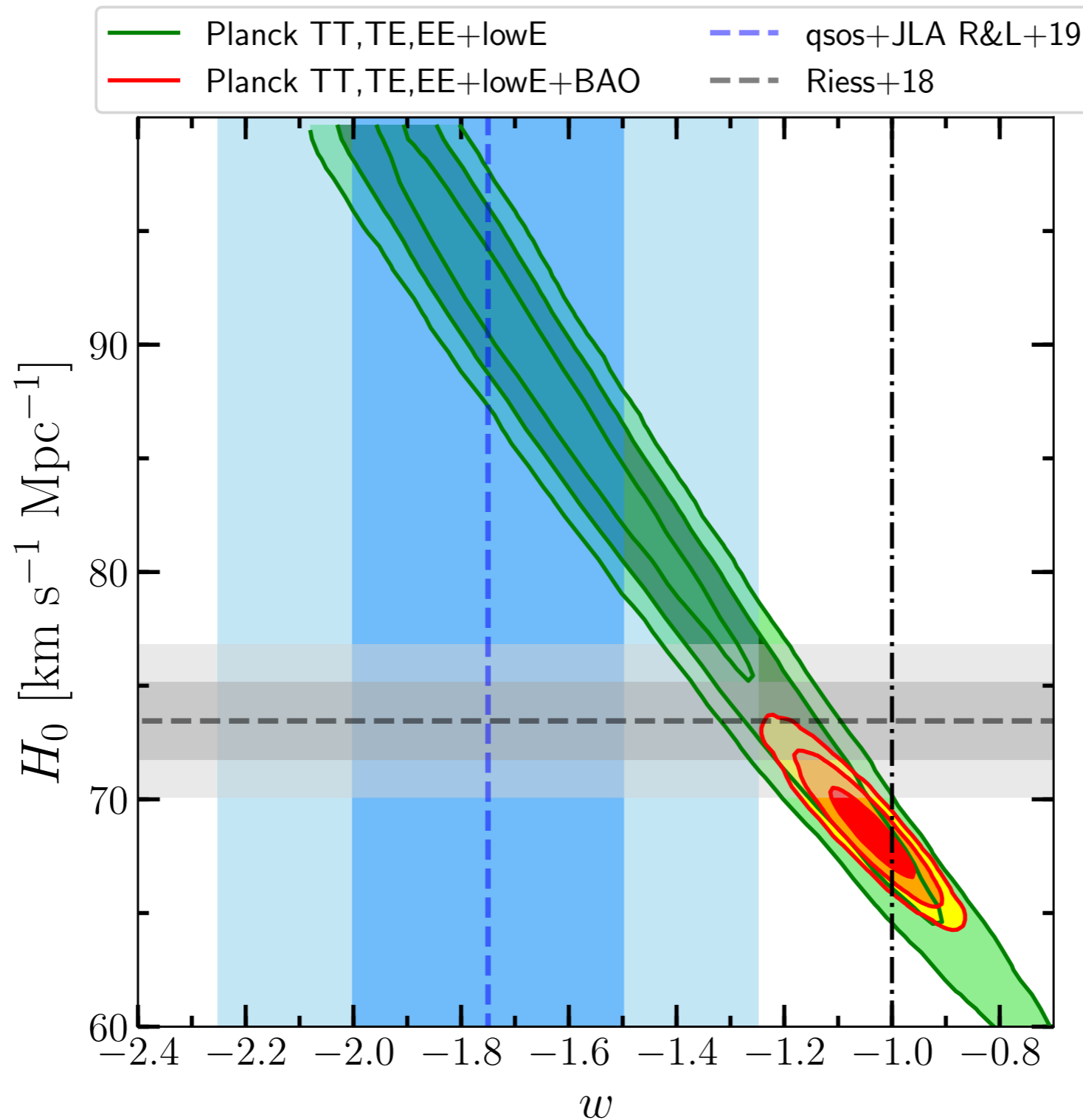
Riess et al. (2019)

$$H_0 = 74.03 \pm 1.42 \text{ km s}^{-1} \text{ Mpc}^{-1}$$



The schism between the early & late Universe

Do we need an extension to the Λ CDM? Maybe yes...



Cosmology with quasars

Test source of systematics

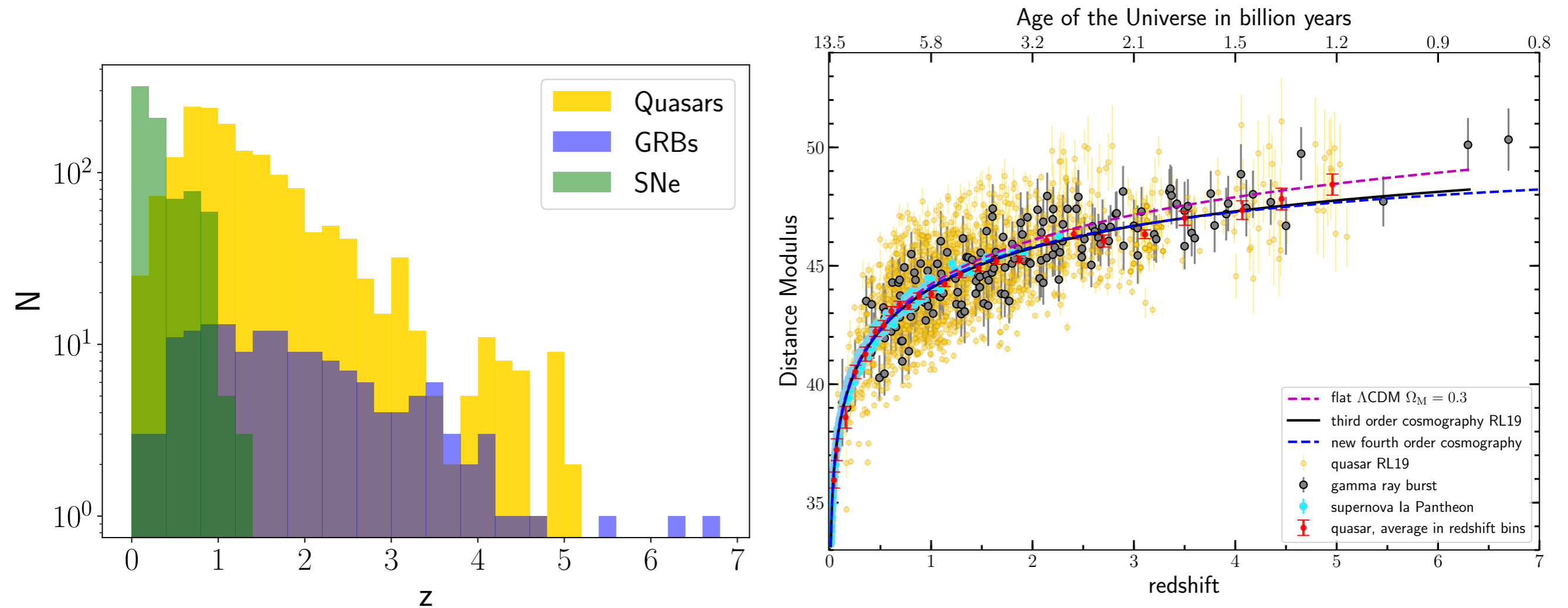
1. Independent samples over a wide redshift range
key to test observed tensions and systematics
2. Different model-independent techniques

The Quasars + SNe + GRBs Hubble Diagram

1598 quasars (Risaliti & Lusso 2019)

1048 Type Ia supernovae - *Pantheon* survey (Scolnic et al. 2018)

160 GRBs (Demianski et al. 2017)

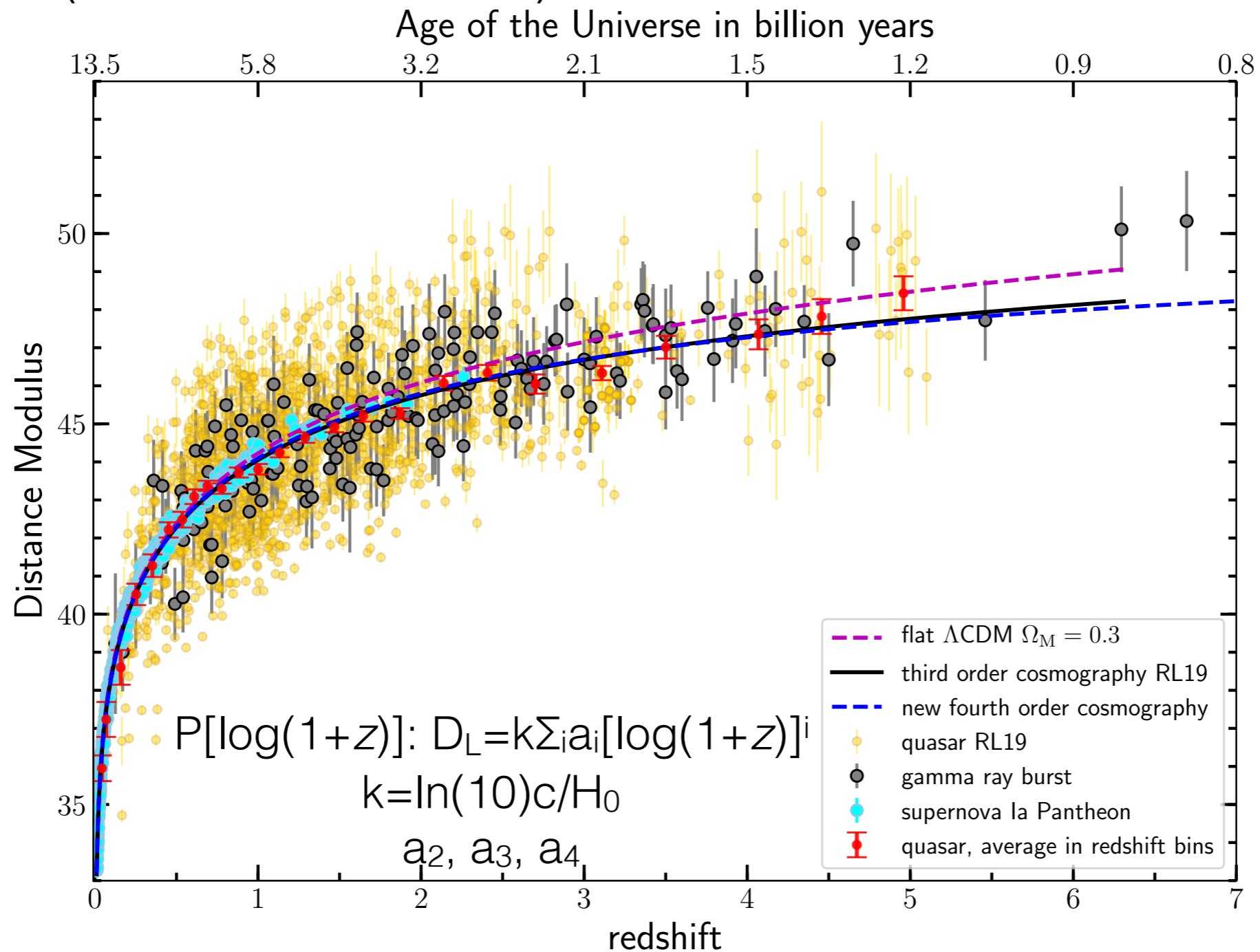


The Quasars + SNe + GRBs Hubble Diagram

1598 quasars (Risaliti & Lusso 2019)

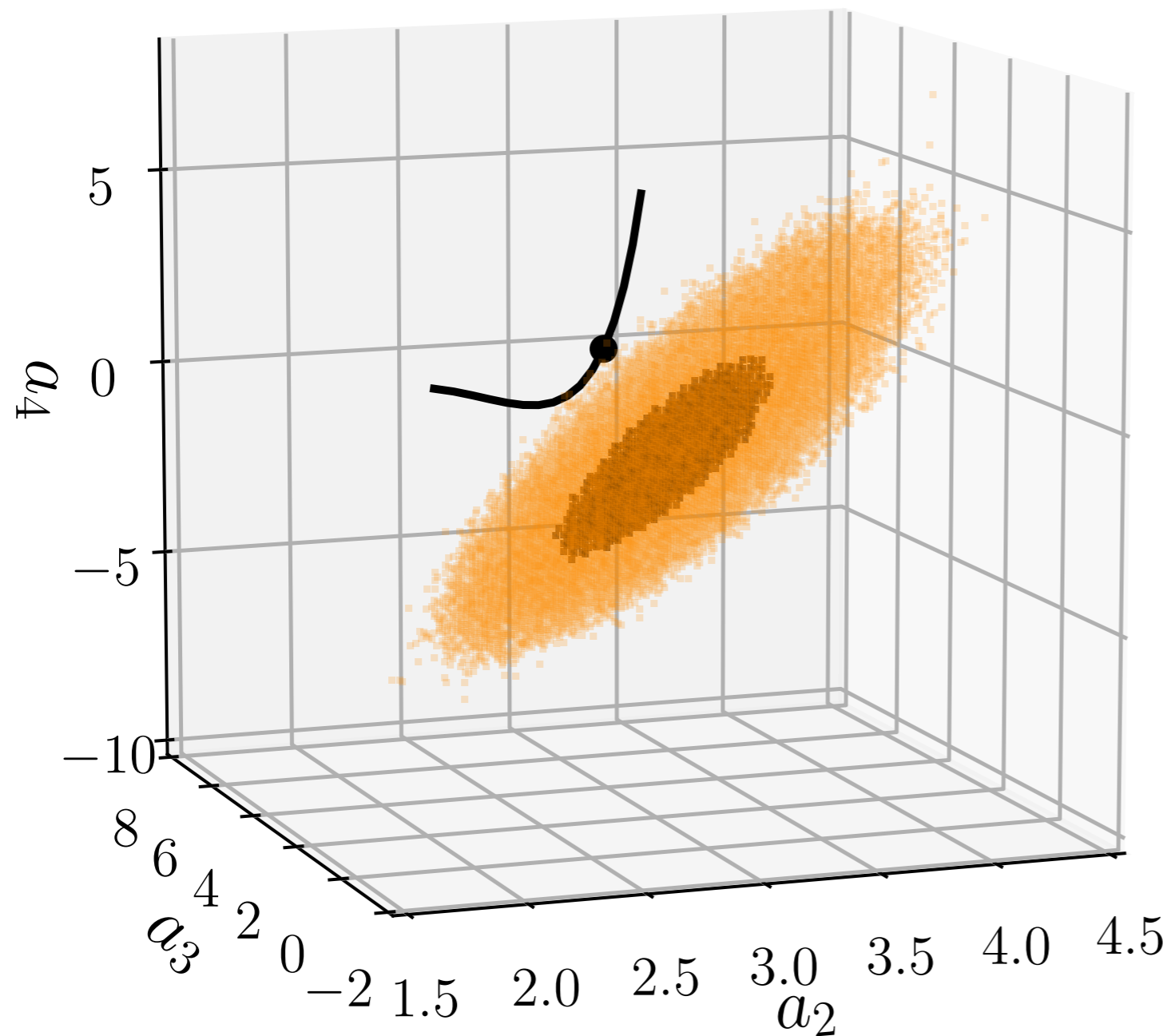
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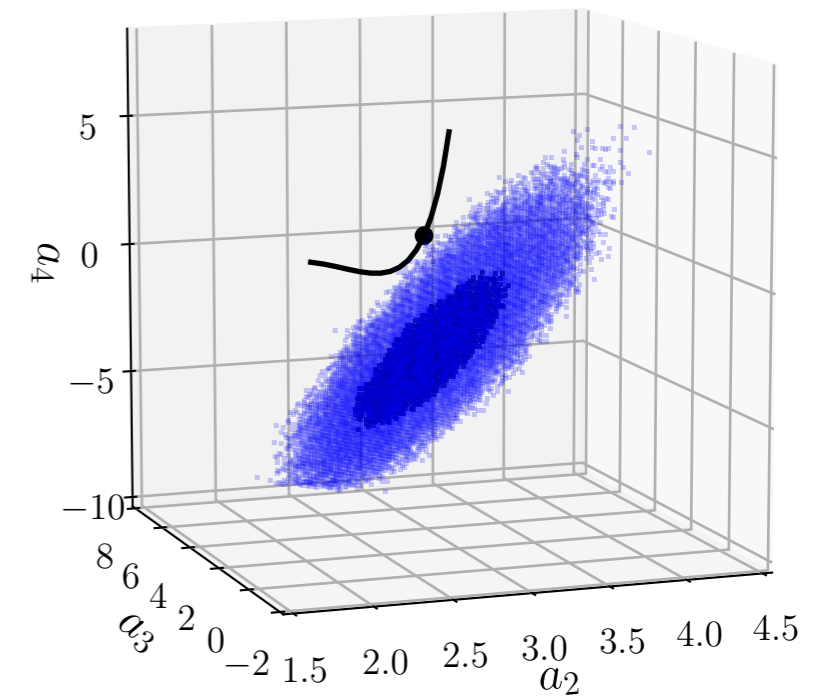


The Quasars + SNe + GRBs Hubble Diagram

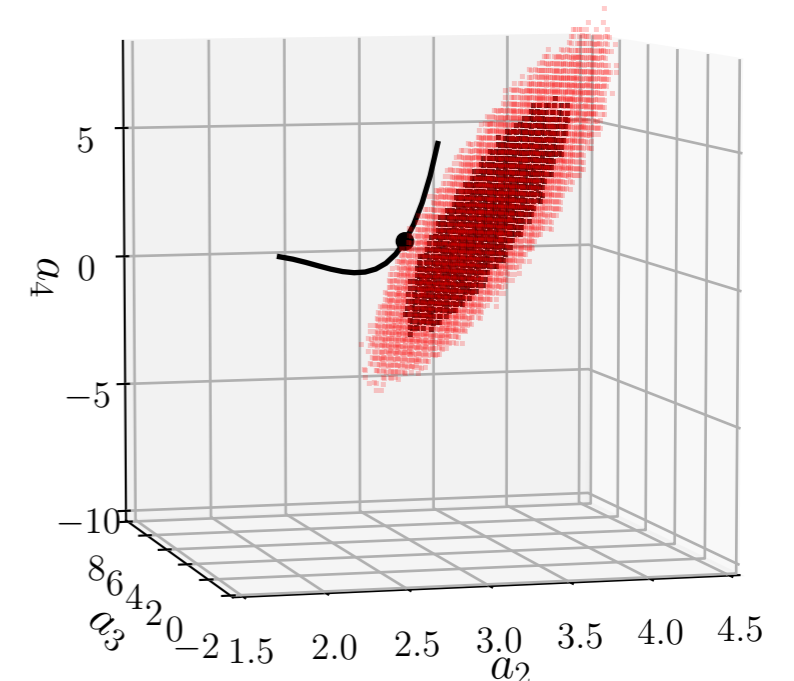
Pantheon, quasars and GRBs



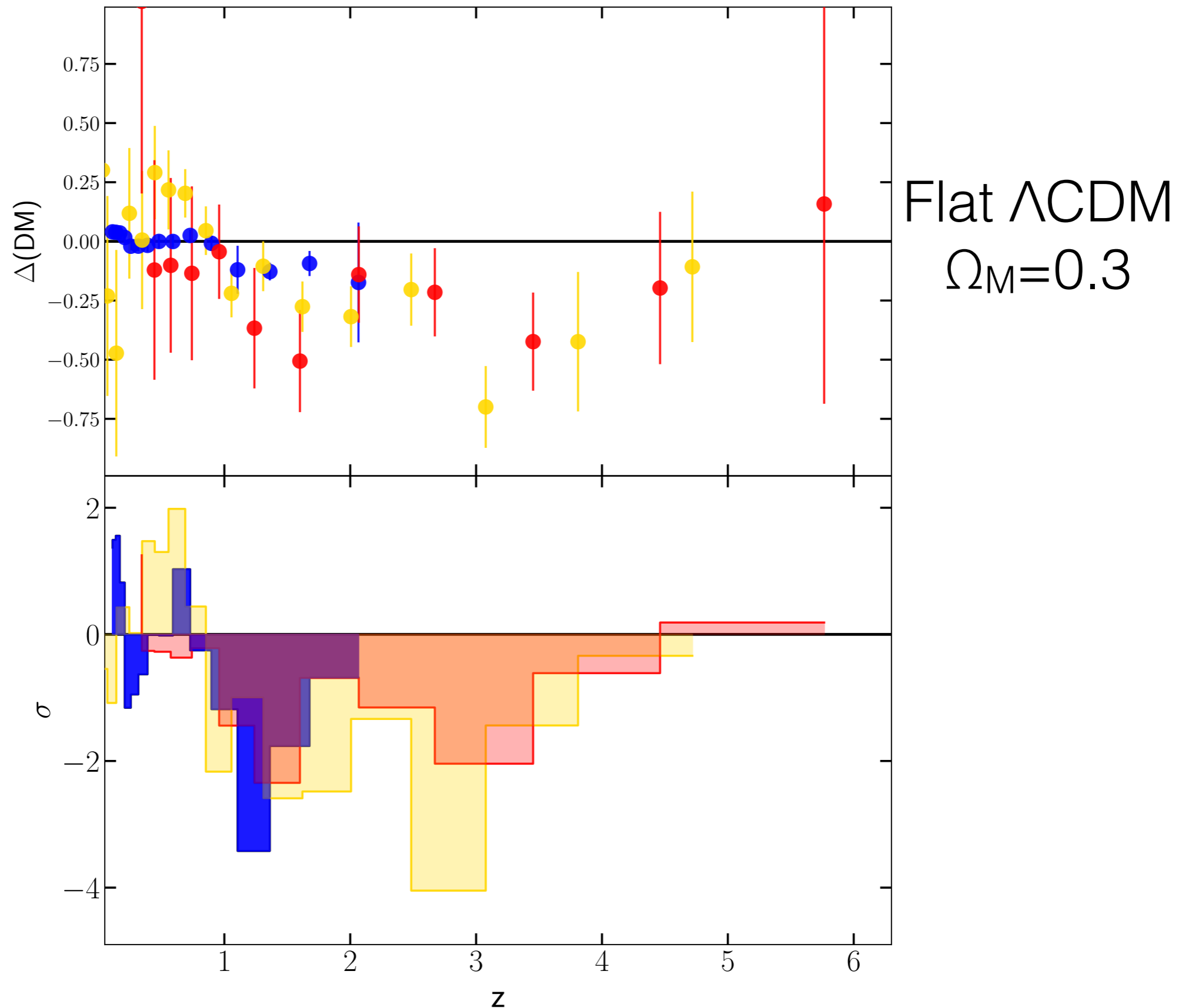
Pantheon and quasars



Pantheon and GRBs



The Quasars + SNe + GRBs Hubble Diagram



To summarise

1. Risaliti & Lusso+19: 4σ tension with the flat Λ CDM model through a model-independent parametrization of a Hubble Diagram of SNe Ia (JLA) and quasars.
2. Confirm the tension with the flat Λ CDM model from a high redshift Hubble Diagram of SNe Ia (Pantheon), quasars and gamma-ray bursts with a statistical significance of $>4\sigma$ (Lusso et al. to be subm).
3. We also confirm that this tension becomes statistically significant (above 3σ) only at high redshifts ($z > 1$) for SNe Ia and quasars taken independently and $\sim 2\sigma$ for GRBs alone (Lusso et al. to be subm).

The completely independent high-redshift Hubble diagrams are fully consistent with each other, strongly suggesting that the deviation from the standard model is not due to unknown systematic effects but to new physics.