



The Dark (and Unknown) Universe

Dark Matter & Dark Energy

Manolis Plionis

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The 2 Main Questions of Contemporary Cosmology: Dark Matter & Dark Energy

95% of the Energy-Density in the Universe is of an unknown form !



The Dark Universe 1: Dark Matter

It is an unknown type of matter which interacts gravitationally with all types of matter, but it does not emit electromagnetic radiation. We know of its existence from the gravitational effects that it has on its surroundings. It has been measured with a variety of independent methods, while its existence is also necessary in order to form cosmic structures without violating the observed amplitude of the CMB temperature fluctuations $<\delta T/T > ~ 10^{-5}$.

The Dark Universe 1: Dark Matter

Astrophysical Evidence for Dark Matter:

- Spiral Galaxies (Rotation Curves)
- Clusters of Galaxies (Virial Theorem, X-ray Emission, Gravitational Lensing)

Cosmological Evidence for Dark Matter:

- Structure Formation
- Primordial Nucleosynthesis (BBPN)
- Join BBPN & Cluster DM
- Cluster M/L vs Global M/L
- SNIa Hubble function + CMB T Fluctuations



Astrophysical Indications for the Existence of Dark Matter: Rotation Curves of Spirals

~1970: Vera Rubin, Ken Freeman measured the rotational velocity of Spiral Galaxies using the 21cm radiation of atomic Hydrogen





Astrophysical Indications for the Existence of Dark Matter: Rotation Curves of Spirals

Spiral Galaxy NGC 4414

NGC 4414 - another spiral galaxy

Spiral Galaxy NGC 4622







NASA and The Hubble Heritage Team (STScl/AURA) • Hubble Space Telescope WFPC2 • STScl-PRC02-03

Astrophysical Indications for the Existence of Dark Matter: Rotation Curves of Spirals What do we expect classically for the rotation curve. From Dynamical arguments we have: $\frac{GM(r)m}{r^2} = \frac{mv(r)^2}{r} \Rightarrow v(r) \propto r^{-1/2}(Keplerian)$

$$instead \Rightarrow v(r) = const. \Rightarrow M(r) \propto r$$



Astrophysical Indications for the Existence of Dark Matter: Clusters of Galaxies

WHAT ARE CLUSTERS OF GALAXIES: THE LARGEST SELF-GRAVITATING, BOUND, MULTICOMPONENT SYSTEMS WITH 100s OF GALAXIES, HOT X-ray EMITTING ICM, RADIO-PLASMAS AND DARK MATTER, WHICH EVOLVE IN A TIGHTLY COUPLED MANNER. DM IS DYNAMICALLY DOMINANT COMPONENT!



Astrophysical Indications for the Existence of Dark Matter: Clusters of Galaxies

First such indication came from the study of the motions of galaxies in the Coma cluster, back in 1933 by Fritz Zwicky. He found that the galaxies mean velocity were larger than their escape velocity, that implied that the Coma cluster should not have existed as such.





Mass Measurement: 2. X-ray Emission

Cluster gas thermalizes to >10⁶ K making clusters strong X-ray sources

Bremsstrahlung radiation: The hot, thin (transparent to its own radiation) plasma emits electromagnetic radiation due to etrajectory change and acceleration when passing near a positive ion



Line emission due to the fact that metals are not completely ionised and so when a fast e⁻ strikes ions with bound e⁻ it causes transition to higher energy level. Excitation lasts little and ions decay to ground state with emission of photon with characteristic energy δE .



Mass Measurement: 2. X-ray Emission



Observations based on space X-ray Observatories (currently XMM & Chandra)



Mass Measurement: 3. Gravitational Lensing

- Based on GR: Radiation is deflected in gravitational fields
- Just as for conventional lenses, images will form at extrema in the light travel time surface (Fermat).
- Normally there will be just one deflected image but, for sufficiently deep gravitational potentials, multiple images form.



Mass Measurement: 3. Strong Gravitational Lensing

observer

O

From simple geometry we get Lens Equation: β =

$$= \theta - \alpha \frac{D_{\rm ds}}{D_{\rm s}}$$

 β = (true) source position θ =(seeming) image position α =(scaled) deflection

Q

$$H(z) = H_0 \left[\Omega_m (1+z)^3 + \Omega_Q \exp\left(3\int_0^z \frac{1+\mathbf{w}(x)}{1+x} dx\right) \right]^{\frac{1}{2}} ,$$

 $D_{\theta} = \frac{1}{1+z} \int_0^z \frac{c}{H(z)} \mathrm{d}z$

For a point mass lens the deflection, α , is given by a simple equation:

$$\alpha = \frac{4GM}{c^2\xi}$$

Images are formed at angles where the deflection eq. and the lens eq. are satisfied simultaneously.

Mass Measurement: 3. Gravitational Strong Lensing

In the case of a well aligned lens-source configuration ($\beta \sim 0$) the deflection, α , is $\alpha \sim \theta$, (the so-called Einstein Ring), which implies that combining the Lens and deflection equations, we have:



$$\beta = \theta - \frac{4GM(\theta)}{c^2\theta} \frac{D_{\rm ds}}{D_{\rm s}D_{\rm d}} \quad \xi = \theta D$$

d

$$\beta = 0 \rightarrow \theta_E = \left[\frac{4GM(\theta)}{c^2} \frac{D_{\rm ds}}{D_{\rm s}D_{\rm d}}\right]^{1/2}$$

Therefore, measuring θ_E observationally, we can estimate the mass of the lens (cluster of galaxy in this case) if we know angular diameter distances of source and lens (which of course depends on the cosmological model)



First Einstein ring observed by VLA Quasar lensed by galaxy

upper: Hubble Iower: Merlin



Mass Measurement: 3. Gravitational Weak Lensing

In this case the distortions of background sources are much smaller and can only be detected by analyzing large numbers of sources to find coherent distortions of only a few percent. The lensing shows up statistically as a preferred stretching of the background objects **perpendicular to the direction to the center of the lens**. By measuring the shapes and orientations of large numbers of distant galaxies, their orientations can be averaged to measure the <u>shear</u> of the lensing field. Since galaxies are intrinsically elliptical and the weak gravitational lensing signal is small, a very large number of galaxies must be used in these surveys.

Mass Measurement: Comparison

(1) Reasonable agreement of Cluster Masses between methods. (2) Weak evidence for validity of General Relativity on cluster scales

COULD THERE BE A TRADITIONAL EXPLANATION: 1. Could Dark Baryons constitute the Dark Matter?

Suggestions exist that failed stars, hot Jupiters etc that have masses below thermonuclear limit, could constitute the DM, ie., Massive Compact Halo Object: MACHOS

These could be detected via gravitational microlensing by the large number of halo stars.

Large observational programs have failed to provide the necessary numbers (not even close...)

2. Could Modification of Gravity make the trick? Could explain rotation curves but not the whole range of DM indications...

Some Basics of Dynamical Cosmology

Friedman's equation: The basic equation of the dynamical evolution of the Universe

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G\rho}{3} + \frac{\Lambda c^2}{3} + \frac{kc^2}{R^2}$$

$$H^2 = \left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3} \left[\rho_m + \rho_\Lambda + \rho_k\right]$$

$$\rho_{\text{total}} = \left[\rho_m + \rho_\Lambda + \rho_k\right] = \frac{3H^2}{8\pi G}$$

Definition of the parametrized density parameter Ω : It is the fractional contribution to global energy density of the Universe.

$$\Omega_i(a) = \frac{\rho_i}{\rho_{\rm total}} = \frac{8\pi G \rho_i}{3H^2}$$

Important relation among Ω 's

 $\Omega_m + \Omega_\Lambda + \Omega_k = 1$

Cosmological Indications for the Existence of Dark Matter: Structure Formation

Cosmological Indications for the Existence of Dark Matter: Structure Formation

Formal Jeans Theory in an expanding background and in comoving coordinates, Using Continuity Equation, Euler's Equation & Poisson's Equation and Linear Perturbation Theory we obtain the basic differential equation for the growth of density perturbations:

$$\ddot{\delta} + 2H\dot{\delta} = 4\pi G \overline{\rho} \delta$$

$$\ddot{\delta} + 2H\dot{\delta} - \frac{3}{2}\Omega_m H^2 \delta = 0$$

Matter Era ($\Omega_m = 1, H = 2/3t$)

$$\ddot{\delta} + 2H\dot{\delta} - \frac{3}{2}\Omega_m H^2 \delta = \ddot{\delta} + \frac{4}{3t}\dot{\delta} - \frac{2}{3t^2}\delta = 0$$

Solution:
$$\delta(t) \approx A t^{2/3} + B t^{-1}$$

Growing mode and a decaying mode

Cosmological Indications for the Existence of **Dark Matter: Structure Formation**

Density fluctuations in a flat, matter dominated Universe grow as

$$\delta \propto A t^{2/3} \propto R(t) \propto \frac{1}{(1+z)}, \qquad \delta << 1$$

- Baryonic Matter fluctuations can ONLY have grown by a factor $(1+z_{dec}) \sim 1000$ by today
- for $\delta \sim 1$ (just entering non-linearity today) require $\delta \sim 0.001$ at recombination
- $\delta \sim 0.001 \implies \delta T/T \sim 0.003$ at recombination But CMB $\rightarrow \delta T/T \sim 10^{-5}$!!!

R (NO DARK MATTER)

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Cosmological Indications for the Existence of Dark Matter: Primordial Nucleosynthesis

In the first three minutes, the Universe was hot enough for nuclear reactions to take place

Protons and neutrons formed ²H (deuterium), ³He and ⁴He. Deuterium is reactive and was used up quickly

⁴He is most stable and, within 3 minutes, made up 25% of the Universe

 $^{2}H + ^{2}H \rightleftharpoons ^{3}H + p$ $^{2}H + ^{2}H \rightleftharpoons ^{3}He + n$ $^{3}H + ^{2}H \rightleftharpoons ^{4}He + n$ $^{3}He + ^{2}H \rightleftharpoons ^{4}He + p$ Because ²H is so reactive, most neutrons in ²H will end up in ⁴He

$$Y = (0.230) + (0.0011) \ln \left(\frac{10^{10} n_{\rm B}}{n_{\gamma}} \right)$$
$$+ (0.013) (N_{\rm L} - 3)$$
$$+ (0.014) (t_{1/2} - 10.6 \,{\rm min})$$

Cosmological Indications for the Existence of Dark Matter: Primordial Nucleosynthesis

From studies of:

 Quasar spectra, we measure absorption lines due to primordial, intervening, gas clouds

~4% of critical density is baryonic

Cosmological Indications for the Existence of Dark Matter: Join Cluster DM & BBPN

BASIC HYPOTHESIS is that the DM and Baryonic mix in Clusters of galaxies corresponds to the Universal value.

Then compare estimate of Total Cluster Mass with baryonic mass (galaxies and gas) to BBPN value to get Ω_{b} .

First application by White et al. 1993, Nature: "The baryon content of galaxy clusters - A challenge to cosmological orthodoxy".

$$\frac{M_b}{M_{Total}} = \frac{\Omega_b}{\Omega_m} \approx 0.15 \Longrightarrow \Omega_m \approx \frac{0.04}{0.15} \approx 0.27$$

23% of total mass-energy density is DARK MATTER

Cosmological Indications for the Existence of Dark Matter: Measuring the values of Ω_i

JUST A FEW BASIC METHODS (many more...)

M/L & baryonic matter in Clusters of Galaxies → Ω_m
 Temperature fluctuations of the CMB →Ω_κ
 Hubble diagram (SN Ia) → Ω_m + Ω_Λ

Cosmological Indications for the Existence of Dark Matter: Ω_m from Cluster M/L vs Universal Value

Estimate Cluster MASS to LIGHT ratio and compare with GLOBAL VALUE, assuming that mean $(M/L)_{clusters} = (M/L)_{global}$ on average

- A. Determine Mass of clusters
- B. Determine luminosity of clusters
- C. Determine Global M/L of Universe

$$\frac{M}{L} \approx 340h \frac{M_{\oplus}}{L_{\oplus}}$$

$$\Phi(L) = C \left(\frac{L}{L^*}\right)^a \exp(-L/L^*)$$
$$< L >= \int L \Phi(L) dL \propto L^* \Gamma(a+2)$$

$$\left(\frac{M}{L}\right) = \frac{\rho_0}{\langle L \rangle} = \frac{\rho_{crit}\Omega_m}{\langle L \rangle} \approx 1520\Omega_m h \frac{M_{\oplus}}{L_{\oplus}}$$

Preliminary Result: $\Omega_m \approx 0.23$

Cosmological Indications for the Existence of Dark Matter: Ω_k from $\delta T/T$

Different angular size on the last scattering surface of the same characteristic scale for different geometries !

• Spherical Harmonic expansion \rightarrow

 $\Delta T/T(\theta, \varphi) = \sum a_{lm} \Upsilon_{lm}(\theta, \varphi).$

netries! ,φ).

Then the Power–Spectrum ($C_{l} = \langle |a_{lm}|^2 \rangle$) represents the amplitude of the contribution of fluctuations from different angular scales (note that $\theta = \pi/l$).

• Fluctuations at θ < 2° are due to the oscillations of the photo-baryonic fluid before recombination

Fluctuations at the scale of the acoustic horizon (at recombination) create a very strong peak in the power-spectrum at : $/ \sim 220/(1-\Omega_k)^{1/2}$

Cosmological Indications for the Existence of Dark Matter: $\Omega_m + \Omega_{\Lambda}$ from SNIa Hubble function

Supernova Ia's appear dimmer than what expected in decelerating Universes, therefore they are further away and this implies accelerated expansion of the Universe $\Rightarrow \Omega_{\Lambda} > 0$ and more detailed theoretical calculation $\Omega_{\Lambda} > 1/3$.

Nature of the dark matter—Hot or cold

- Hot dark matter is relativistic at the collapse epoch and free-streaming out the galaxy-sized over density. Larger structure forms early and fragments to smaller ones.
- Cold DM is non-relativistic at de-coupling, forms structure in a hierarchical, bottom-up scenario.
 HDM is tightly bound from observation and LSS formation theory

Conclusions

 Dark Matter, of a Coldish type which selfinteracts gravitationally, definitely exists !
 It constitutes ~23% of total mass-energy density + 4% baryons + 73% DE.... !
 There are many candidates for DM particles:

Axions, Neutralinos, Gravitinos, Axinos, Kaluza-Klein Photons, Kaluza-Klein Neutrinos, Heavy Fourth Generation Neutrinos, Mirror Photons, Mirror Nuclei, Stable States in Little Higgs Theories, WIMPzillas, Cryptons, Sterile Neutrinos, Sneutrinos, Light Scalars, Q-Balls, D-Matter, Brane World Dark Matter, Primordial Black Holes, ...